# EXO-200: Results, Plans for Phase-II, and Beyond

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# Origin of Neutrino mass

- From neutrino oscillation experiments we know that neutrino has a non-zero mass. We have measured relative mass-squared differences of different states
  - But what about **absolute mass values**?
  - How do masses of the different states align (*normal, inverted* hierarchy)?
  - What is the origin of the mass term (*Dirac, Majorana*)?
- One of the most promising ways to answer these questions is to study a peculiar type of radioactive decay, forbidden in the SM!

Composition	Elementary particle
Statistics	Fermionic
Generation	First, second and third
Interactions	Weak interaction and gravitation
Symbol	$v_e, v_\mu, v_\tau, \overline{v}_e, \overline{v}_\mu, \overline{v}_\tau$
Antiparticle	Antineutrinos are possibly identical to the neutrino (see <i>Majorana fermion</i> ).
Theorized	v <sub>e</sub> (Electron neutrino): Wolfgang Pauli (1930)
	$v_{\mu}$ (Muon neutrino): Late 1940s
	v <sub>r</sub> (Tau neutrino): Mid 1970s
Discovered	v <sub>e</sub> : Clyde Cowan, Frederick Reines (1956)
	v <sub>µ</sub> : Leon Lederman, Melvin Schwartz and Jack Steinberger (1962)
	v <sub>T</sub> : DONUT collaboration (2000)
Types	3 – electron neutrino, muon neutrino and tau neutrino
Mass	Small, but non-zero. See the mass section.
Electric charge	0 e
Spin	1/2
Weak hypercharge	-1
B-L	-1
x	-3
	Composition Statistics Generation Interactions Symbol Antiparticle Theorized Discovered Discovered Mass Electric charge Spin Weak hypercharge B - L

From Wikipedia

### Double beta decay

- Two-neutrino mode is a Standard model process observed for several isotopes, but is extremely rare
- *Neutrinoless* mode violates lepton number conservation
  - can only happen if neutrinos are massive Majorana particles
  - provides information about absolute mass scale
  - has never been observed\*
- Main goal of EXO-200 is to search for the neutrinoless mode



\* a controversial discovery claim exists by a sub-group of Heidelberg-Moscow collaboration [H.V. Klapdor-Kleingrothaus and I.V. Krivosheina Mod. Phys. Lett., A21 (2006) 1547]

#### Detecting double-beta decay



Illustration from P.Vogel, arXiv:hep-ph/0611243, Assumes 2% resolution and 1e2 (1e6 in insert) ratio of 2nu/0nu

- In the two-neutrino mode electrons have to share energy with undetectable neutrinos
  - A calculable, but broad and featureless spectrum
- In the neutrinoless mode, a monoenergetic peak is expected at Qvalue
  - Good energy resolution is essential
  - Large Q-value is preferred

#### EXO-200

- ~200 kg of Xe enriched to 80.6% in  $^{136}$ Xe
  - ~175 kg in liquid phase inside a cylindrical Time Projection Chamber
  - ~100 kg current fiducial mass
- Located at 1585 m.w.e. in the Waste Isolation Plant near Carlsbad, NM
  - Muon rate reduced to the order of  $10^{-7}$  Hz /cm<sup>2</sup> /sr
  - Salt has inherently lower levels of U/Th (<100 ppb), compared to rock
  - Low levels of Rn (~20 Bq/m<sup>3</sup>)
- Carefully selected radioactively clean materials, rigorous cleaning procedures, detector installation inside class 1000 clean room
  - Goal of **40 counts/2 yrs** in 2σ 0nu energy window (assuming 140 kg LXe, 1.6% resolution)
  - M. Auger et al., JINST 7 (2012) P05010 and D.S. Leonard et al., NIM A 591 (2008) 490

## EXO-200 time projection chamber (TPC)







### Single- vs. Multi-site events



Very useful in identifying gamma backgrounds!

#### Two-neutrino result



EXO-200 **discovered** the 2v mode in <sup>136</sup>Xe and provided the **most accurate** measurement of a 2v half-life among all isotopes

Chicago, 01/2016

## Measurement of two-neutrino mode half-life

 2v measurement is not limited by statistics, so we used a strict fiducial cut to leave only the best-modeled core volume – 66.2 kg of <sup>136</sup>Xe

Component	Error (%)
Normalization	2.6
SS fraction	0.77
Backgrounds	1.3
Statistical	0.76
Total	2.83



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Rate of 2v decay vs. apothem

#### Dataset for neutrinoless analysis



- **477.60±0.01** live days collected between 10/2011 and 09/2013
- Larger fiducial volume to maximize exposure
- **100.0±3.4** kg·yr <sup>136</sup>Xe

## Systematics budget

• Signal detection efficiency:

Source:	Signal efficiency [%]:	Relative error [%]:
Summary from 2-nu	93.1	0.9
Partial reconstruction	90.9	7.8
Fiducial volume		3.4
Total:	84.6	8.6

• Region-of-interest (ROI) backgrounds:

Source:	Relative error [%]:
Background shape distortion	9.2
Choice of background model components	5.7
Variation of energy resolution over time	1.5
Total:	10.9

- Deviation between  $\beta$  and  $\gamma$  energy scale ( $E_{\beta} = B \cdot E_{\gamma}$ ): **B** = 0.999 ±0.002
- Single-site fraction error\*: (Data MC)/Data = 9.6%

#### Final fit



#### Single-site zoom-in



#### **2σ ROI breakdown for major backgrounds**

	Events	BI, 1e-3 /kg/yr/keV*	*normalized to total exposure
Th-232	16.0		(124 kg*yr).
U-238	8.1		<sup>136</sup> Xe exposure is 100 kg·yr
Xe-137	7.0		
Total	_31.1±1.89(stat)±3.3(syst)	1.7±0.2	
Compare to p.			

#### 90% C.L. Limit



- The event excess is compatible with background fluctuation at ~1.2σ
- Median 90% C.L. U.L. limit assuming adequate "background-only" model: >1.9e25 yrs
- Limit from the fit to the actual data:
  >1.1e25 yrs
- ~1.5 " $\sigma$ " away from the median
  - 14% of signal-free toys have limits worse than this

#### Context



# Other BSM physics: 0v Majoron-mediated decays

- $(A, Z) \rightarrow (A, Z+2)+2e^{-} + \chi_{0}$   $(A, Z) \rightarrow (A, Z+2)+2e^{-} + 2\chi_{0}$ Massless or light boson ("Majoron")
- Originally described as Goldstone boson of lepton number symmetry breaking
- Possible dark matter candidates; speculated to be involved in other astrophysical processes
- Characteristic spectral shape that can be searched for in EXO-200
- No stat. significant evidence was found. Limits on coupling constants among strongest to date
- PRD 90, 092004 (2014)



## Other SM physics: 2v decay to excited states

- The two-neutrino mode of double beta decay is allowed in the SM
- The decay may also proceed to the first excited state (01<sup>+</sup>), if energetically accessible
- It is interesting to try and measure this process, because it may provide additional constraints on NME
  - Some NME uncertainties are common for both decays, so cancelations are possible, potentially leading to more accurate extraction of Majorana mass from neutrinoless decay
- It may also test exotic alternative mechanisms for double beta decays



#### Other SM physics: 2v decay to excited states

- Decay to 2<sup>+</sup><sub>1</sub> state from the ground state of <sup>136</sup>Xe has larger Q-value, but is highly suppressed due to angular momentum
- The primary interest for us is thus decay to 0<sub>1</sub><sup>+</sup> state, which results in subsequent emission of 760.5 keV and 818.5 keV gammas
- The expected rate is calculated to be smaller than to the ground state by a factor of 3915 due to PSF, and by additional factor of 2.9 due to NME
  - Expected half-life is thus 2.5e25 yrs, higher than expected sensitivity of EXO-200
  - But NME uncertainties could lead to overestimation, as was the case of <sup>100</sup>Mo and <sup>150</sup>Nd



# Final discriminator. Validating agreement of training and real events



#### Data and MC event distribution of discriminator variable for $2\nu\beta\beta$ decays to ground state

Data and MC event distribution of discriminator

variable for <sup>228</sup>Th calibration source

## Sensitivity improvement



- Using the BDT-based discriminator variable increases median sensitivity by a factor of 3-4, compared to the usual EXO analysis approach
- Expected median sensitivity: T<sub>1/2</sub> > 1.7e24 yr @90% C.L.

## Final fit

- 2-D ML fit in Energy and Discriminator observables
- Same dataset as in the last neutrinoless analysis (100 kg·yr <sup>136</sup>Xe exposure)
- Same background components, plus excited decay signal component



Final fit results. Projection on Discriminator variable

#### Result

- Best-fit value of 43 events
- Statistically insignificant
  - Consistent with NULL at 1.6  $\sigma$
- Corresponding half-life limit: T<sub>1/2</sub> > 6.9e23 yr @90% C.L.
- Roughly a factor of two weaker than median sensitivity



#### Other recent EXO-200 papers

- J.B. Albert et al. "Measurements of the ion fraction and mobility of alpha and beta decay products in liquid xenon using EXO-200"
  - arXiv:1506.00317 (Jun 2015), submitted to Phys. Rev. C.
- J.B. Albert et al. "Investigation of radioactivity-induced backgrounds in EXO-200"
  - Phys. Rev. C 92 (2015) 015503

#### EXO-200: Plans for Phase II

- Following the WIPP accidents on Feb 2014, access to the underground was discontinued and EXO-200 has stopped collecting data
- The access is largely restored by now and EXO-200 is making preparations to restart data-taking
- In May 2015, a panel of experts reviewed the case for continuation of EXO-200. Following the review, DOE has approved the restart and collection of 3 more years of data in June 2015
- Current plan is to restart by the end of the year

## Expected improvements in Phase-II

#### Analysis

- Reduce <sup>137</sup>Xe background with a new veto cut for events consistent with <sup>136</sup>Xe neutron capture γs. Up to ~40% expected reduction
- Using pulse rise time and U-wire induction signals to improve SS/MS discrimination. Simple 1D cut already moves 40% of γs, while only 15% of signal, from SS to MS. More could be possible with continuous variable

#### Hardware

- Electronics upgrade to reduce noise in the light channel.
   Expected to decrease the energy resolution at Q-value from ~1.5% to ~1%
- "Deradonator" system to flush the air gap between cryostat and lead shield with Rn-suppressed air. Expected to remove or reduce "remote" U background

#### EXO-200 Phase-II Estimated sensitivity



### From EXO-200 to nEXO

- EXO-200 demonstrated principle of a homogenous TPC capable of controlling backgrounds by a *combination* of energy resolution, event topology, and event location
- nEXO will take better advantage of all three (pending certain R&D):
  - Bigger detector w/o central cathode better discrimination of external bkgs with position dependent fit
  - Better photodetection and new charge collection scheme with cold electronics – better energy resolution and multiplicity metrics



#### From EXO-200 to nEXO: bigger, with cleaner core volume



- Based on EXO-200 experience, we plan to do standoff distance fit in (almost) whole volume
- No central cathode means no source of Bi-214 gammas in the core volume



#### The role of standoff in background control of a big detector



Full volume **likelihood fit will always outperform simple fid. cut**, as long as one can model the shape of probability density functions adequately

#### From EXO-200 to nEXO: higher gain photodetectors, bigger coverage



- Combine light/charge for best resolution
- APD noise limits resolution in EXO-200
- With barrel placed SiPMs, assume 1% for nEXO (but even 0.5% not impossible)



#### SiPM technology is almost there!



- Both Hamamatsu and FBK, basically, achieved min.
   PDE@175nm requirement (15 abs.%, w/o wavelength-shifter)
- Other parameters also improve from one production to another
- FBK readily provides bare devices for ultimate radiopurity

For recent results of nEXO R&D effort on SiPMs: *I.Ostrovskiy et al. IEEE TNS* **62** (2015) 1825.

#### From EXO-200 to nEXO: Full list of advancements

What	Why
~30x volume/mass	To give sensitivity to the inverted hierarchy
No cathode in the middle	Larger low background volume/no <sup>214</sup> Bi in the middle
6x HV for the same field	Larger detector and one drift cell
>3x electron lifetime	Larger detector and one drift cell
Better photodetector coverage	Energy resolution
SiPM instead of APDs	Higher gain, lower bias, lighter, E resolution
In LXe electronics	Lower noise, more stable, fewer cables/feedthroughs, E resolution, lower threshold for Compton ID
Lower outgassing components	Longer electron lifetime
Different calibration methods	Very "deep" detector (by design)
Deeper site	Less cosmogenic activation
Larger vessels	5 ton detector and more shielding

#### nEXO R&D is in full swing to address remaining challenges

- High Voltage
  - Need 50 kV to maintain the same field as in EXO-200
  - Most LXe experiments had HV problems
  - Phase 1: <3kg setup confirms breakdown from well polished surfaces at ~300 kV/cm
  - Phase 2: 100kg "miniEXO" test setup in progress. Preliminary indication is that EXO-200 problems are specific to EXO-200
  - Phase 3: Planned full scale nEXO segment with final materials. Designed in coordination with LZ



"mini-EXO" test setup

#### nEXO R&D is in full swing to address remaining challenges

- Cryogenic electronics
  - Cables are substantial contributor to background budget in EXO-200 and nEXO plans to further increase granularity of readout to improve topology discrimination
  - Fully integrated, ultra-low background cold electronics has not been built before
  - nEXO is working on a proof of principle chip for a 10x10 cm<sup>2</sup> tile, to be tested for radiopurity and in performance in LXe



Assumes simple tile charge collection system with interleaved strips and EXO-200 style cables for the remote location cases.

#### nEXO R&D is in full swing to address remaining challenges

- Also working on material radiopurity tests, simulation, calibration ideas
- Limited work on mechanical design of the vessel and cryostat
  - TPC vessel is copper, as in EXO-200
  - Considering carbon-composite cryostat (easier to construct UG, potentially cleaner, would not need as much HFE)
- Cryopit (SNO lab) as primary choice of location
  - <sup>137</sup>Xe background (~25% in EXO-200) is not an issue



## Summary

- EXO-200 has already produced one of the strongest results for neutrinoless, two-neutrino, and Majoron-mediated decays
  - The experiment is currently preparing for Phase-II
  - Phase-II is expected to improve sensitivity to neutrinoless decay by up to a factor of 3, compared to Phase-I (assuming 3 years of data-taking and stated hardware and analysis improvements)
- The next generation experiment, nEXO, is a planned 5-tonne scale-up
  - Because of its multi-parameter capabilities, nEXO has robust discovery potential
  - Its general configuration was validated by successful EXO-200
  - Homogeneity is a desirable feature. Required R&D is in full swing
  - This is a tested collaboration that is known to be capable of successfully executing every phase of an experiment
  - It is essential that this science is done in an effective and timely manner. *nEXO plans* to be ready to start construction project in 2017





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