

# Discovery of the Pion

Anton Kapliy  
April 30, 2008

# Cecil Frank Powell

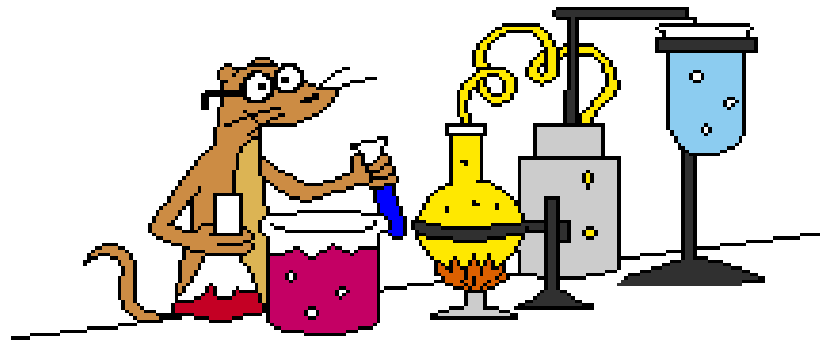
- 1903 - 1969
- Wilson, Rutherford
- University of Bristol
- First pions in 1947
- Nobel in 1950
- Russel-Einstein msto
- CERN



Scanned at the American  
Institute of Physics

# Scientist in the making

- 12 yo — first chemistry book  
*"Found it full of romance"*
- 14 yo — first big experiment
  - generated hydrogen "by the action of granulated dilute sulphuric on granulated zinc"



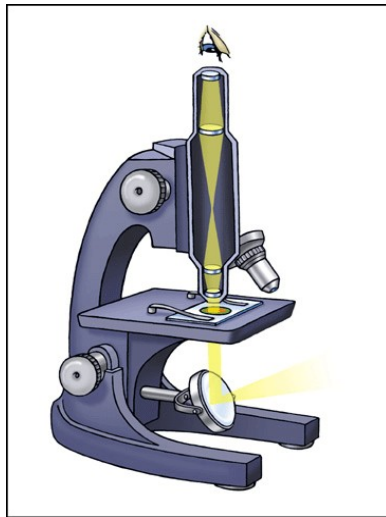
# Early life

- 1920's — Cavendish Labs
- 1932 — wedding, industry temptation
- 1930's-1940's — Bristol, British atomic project
  - Starts experiments with photoemulsions



# Photographic emulsions

- Grains of silver-halide
- Suspended in gelatin
- Ionizing particles create Ag grains
- Develop a photograph



emulsion

glass

50  $\mu\text{m}$

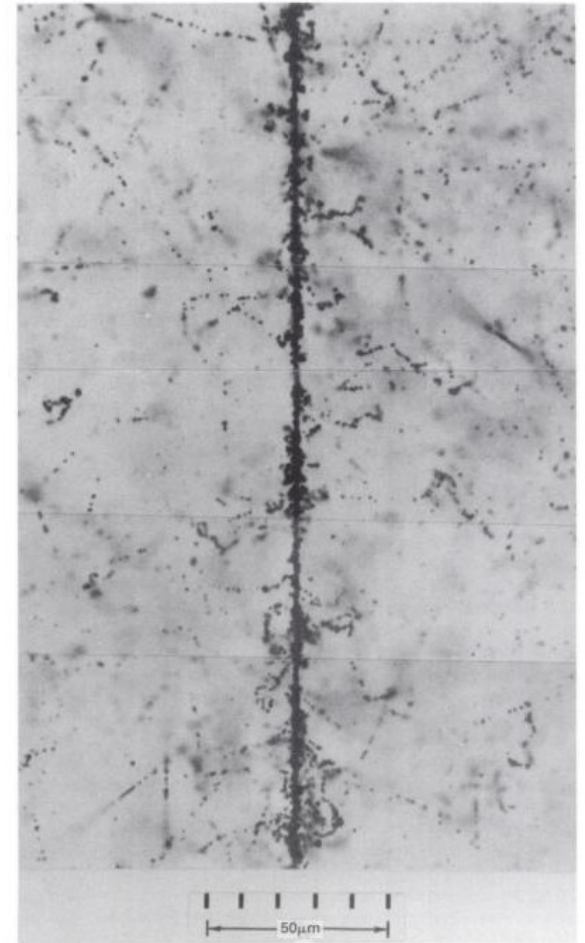


Fig. 5.1 Track of a highly ionizing nucleus in a photographic emulsion. The short tracks produced by  $\delta$  electrons may be clearly seen surrounding the track. Figure courtesy of P. H. Fowler.

# Problems with emulsions

- Low sensitivity to ionization (before 1947)
- Fading over time
- Non-uniformity of active material
- Hard to observe the tracks (too small)

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## Powell's response:

- Collaboration with the industry: Kodak, Ilford
- New sensitizing methods
- State of the art microscopes

# More problems

- Unable to use B-field to bend trajectories
  - Compare with Wilson (aka cloud) chambers!
- Cavendish dis-endorsement:
  - *incapable of "reliable and reproducible precision measurements"*



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**Powell's response:** 20000 stereoscopic photos => 1600 usable traces  
3000 traces on a 3 cm<sup>2</sup> photographic plate

- But photoplates have continuous uptime
- Can deduce a lot about nature of particles
- Powell's determination
  - *"The right man pushing exactly the right experimental technique at the right time"*

Q: Where do we look for Yukawa mesons?  
A: Cosmic rays!



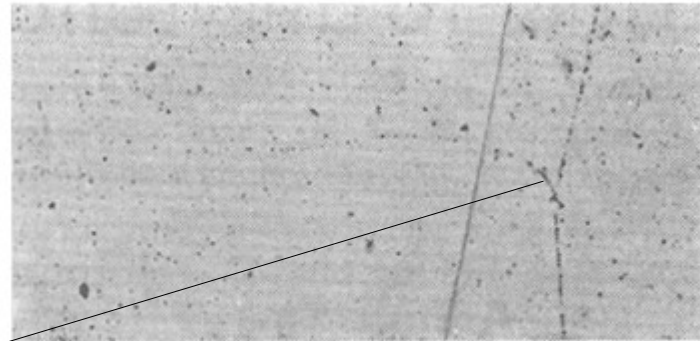
Mt. Pic Du Midi, 10000 ft



Flying balloons, >10000 ft

# First pion

## Nuclear capture of pion



g. 1 a. PHOTOMICROGRAPH OF CENTRE OF STAR, SHOWING TRACK OF MESON PRODUCING DISINTEGRATION. (LEITZ 2 MM. OIL-IMMERSION OBJECTIVE.  $\times 500$ )

- A is the new meson
- B, D, C are likely protons
- Track C goes into the page

Why A is a new meson:  
 electron: range too large  
 proton: scattering too large  
 muon: frequent nuclear interaction

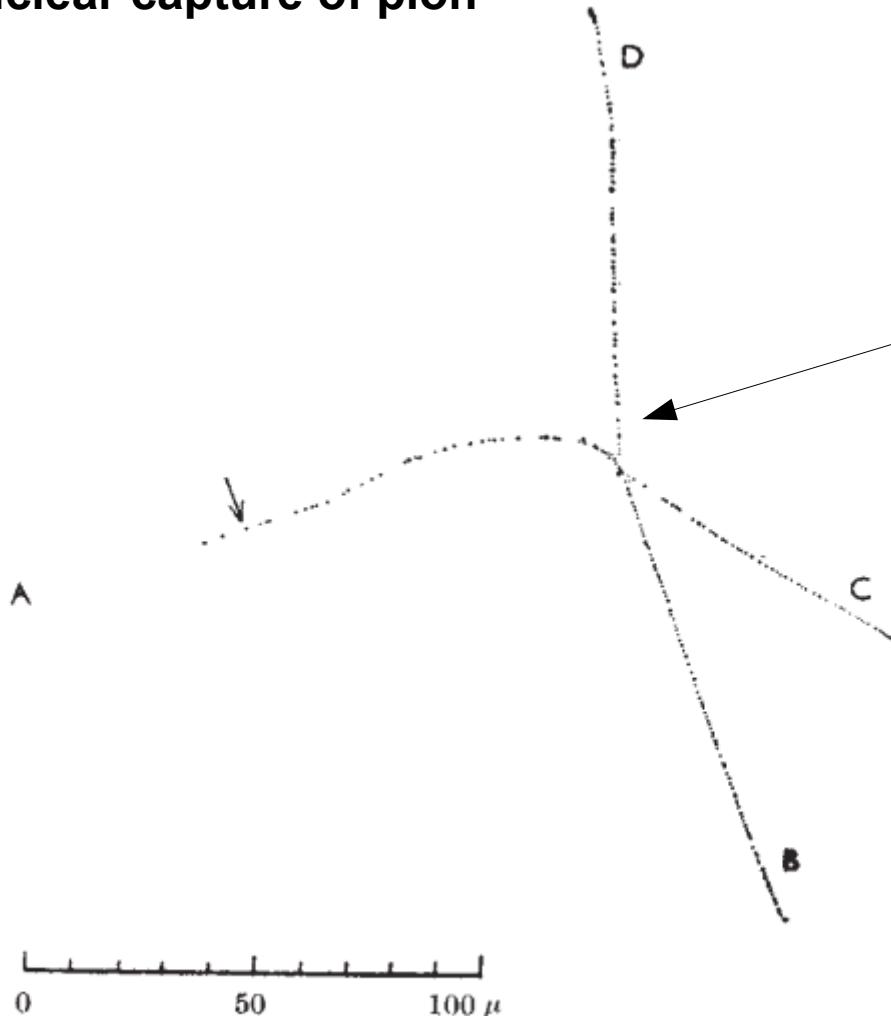
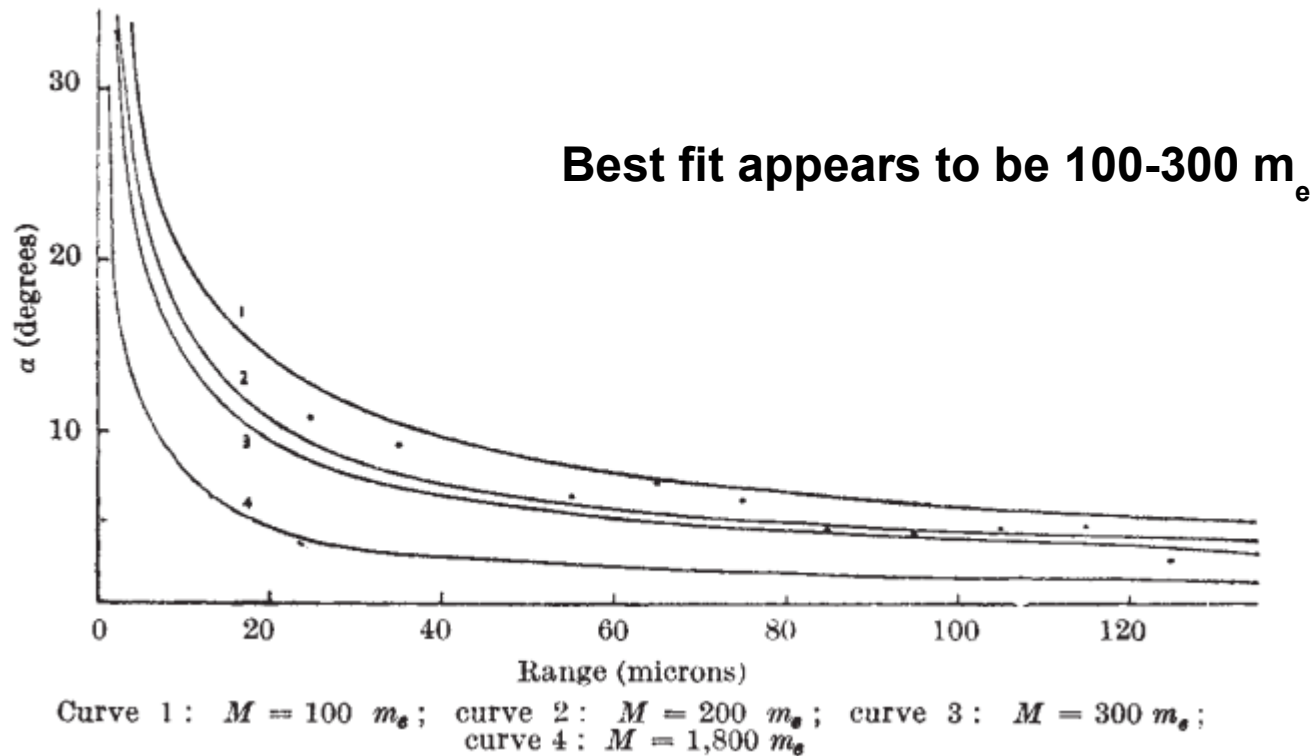


Fig. 1 b. TRACE OF COMPLETE STAR ON SCREEN OF PROJECTION MICROSCOPE, SHOWING PROJECTION OF THE TRACKS IN THE PLANE OF THE EMULSION. TRACK A CANNOT BE TRACED WITH CERTAINTY BEYOND THE ARROW

(Jan 1947, observed by D. Perkins)

# First pion — mass via scattering

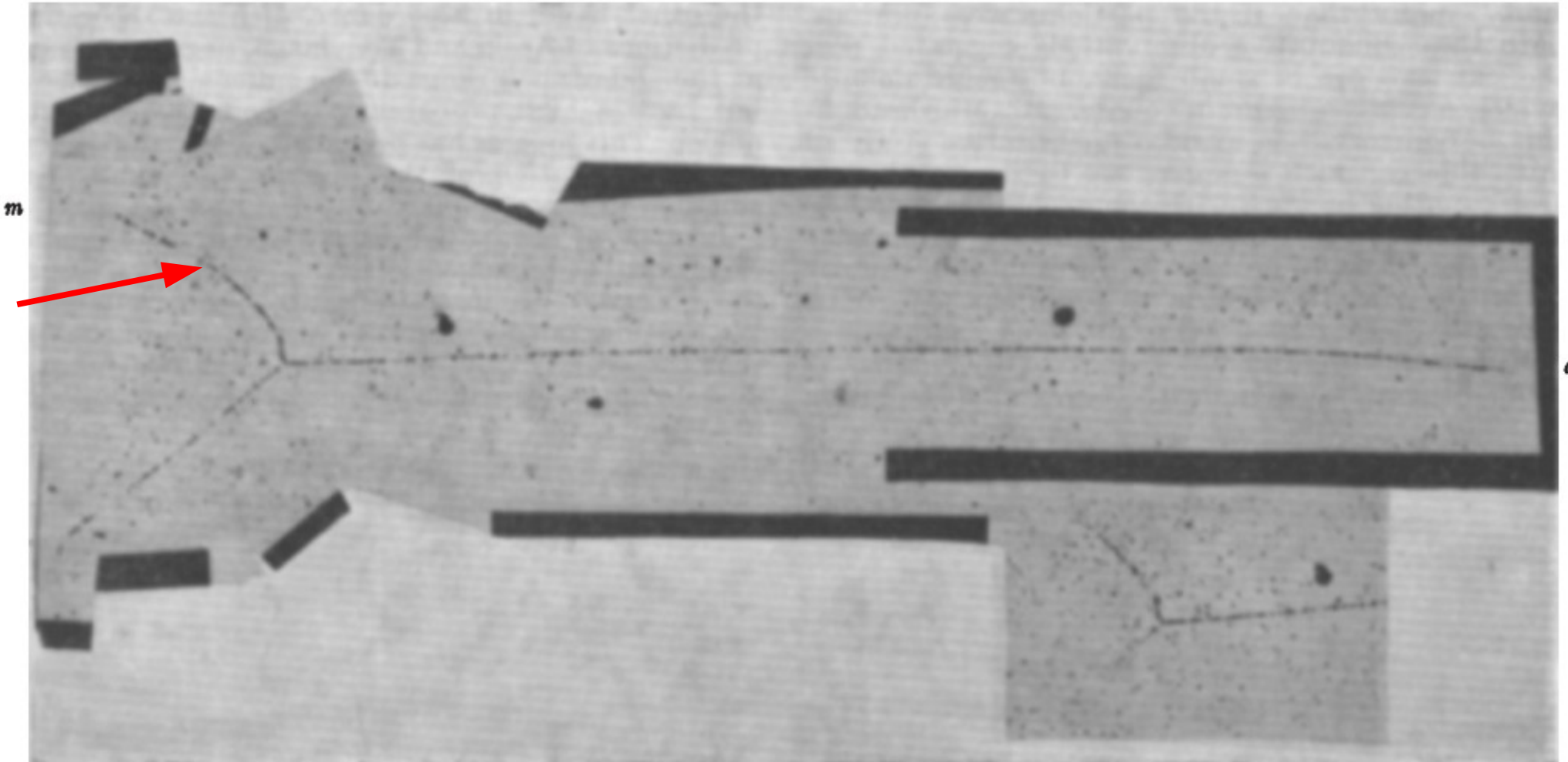


**Horizontal axis:** range of travel in microns

**Vertical axis:** mean scattering angle

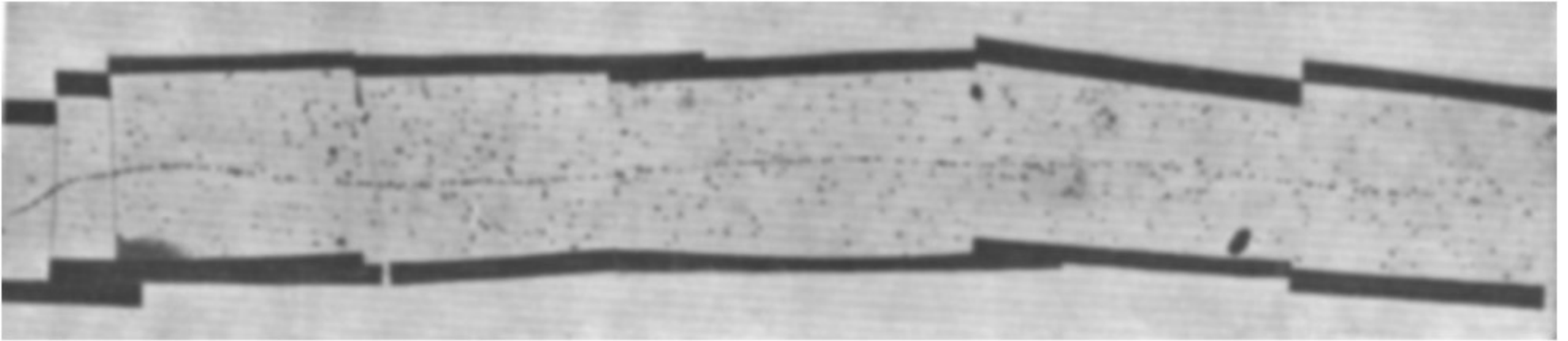
**Theory curves:** Williams's multiple scattering formula

# Another pion capture: $\pi^-$



According to predictions of Araki and Tomonaga, **negative mesons** are usually captured by nucleus at the end of their range

# Event suspected to be $\pi^+$



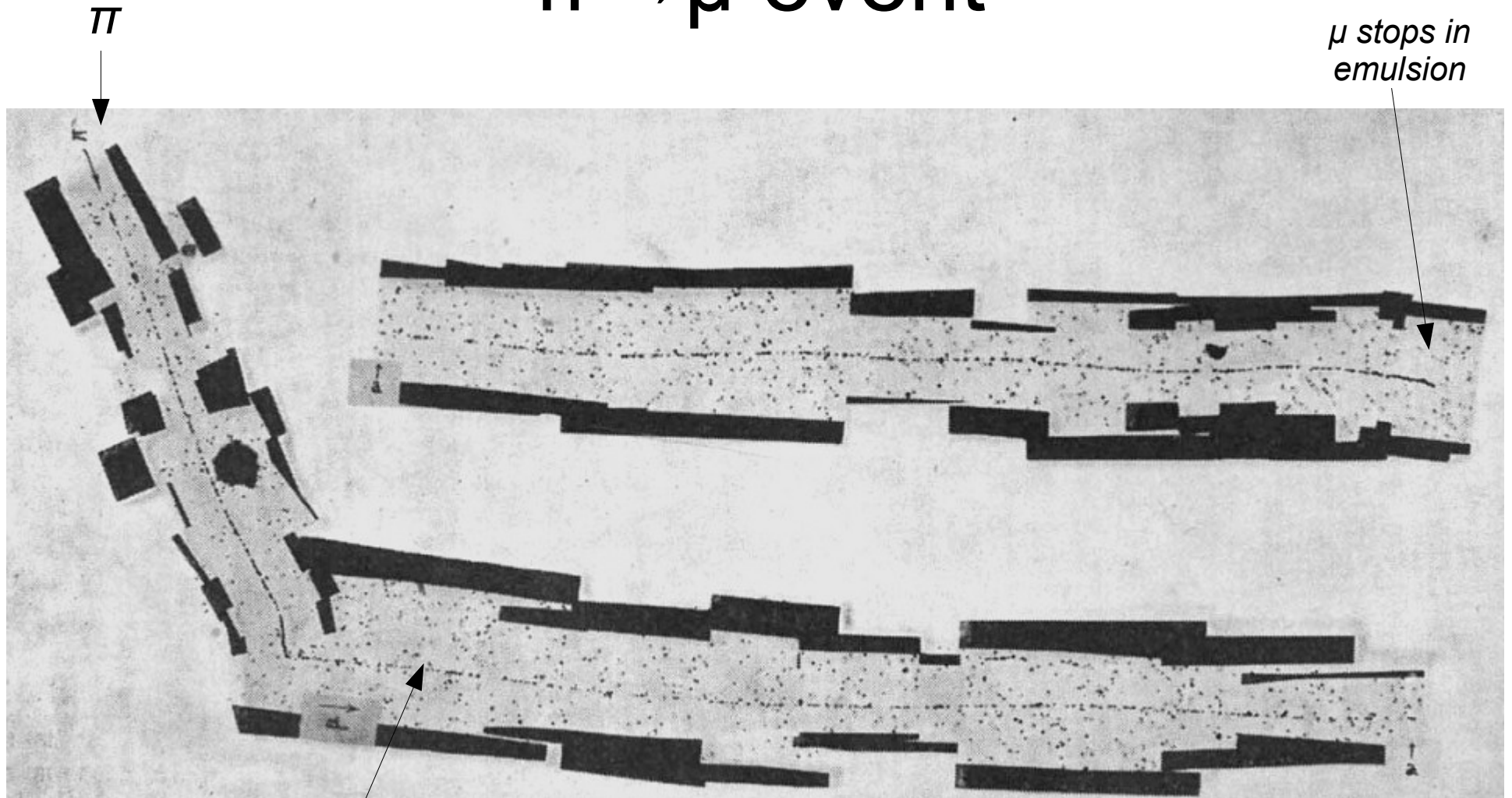
Apparently, no nuclear capture. Comes to rest at the very end of range. A muon probably escaped the emulsion and wasn't detected.

Ionization (distance between grains) suggests a meson of mass **100-230  $m_e$**

Mean scattering angle suggests a meson of mass **250-350  $m_e$**



# $\pi \rightarrow \mu$ event



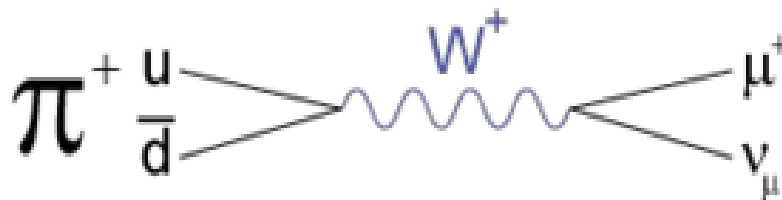
$\mu$  (estimated mass =  $100-300m_e$ )

# Range of muons

Event No.	Range in emulsion in microns of	
	Primary meson	Secondary meson
I	133	613
II	84	565
III	1040	621
IV	133	591
V	117	638
VI	49	595
VII	460	616
VIII	900	610
IX	239	666
X	256	637
XI	81	590

Mean range  $614 \pm 8 \mu$ . Straggling coefficient  $\sqrt{\Sigma \Delta_i^2/n} = 4.3$  per cent, where  $\Delta_i = R_i - \bar{R}$ ,  $R_i$  being the range of a secondary meson, and  $\bar{R}$  the mean value for  $n$  particles of this type.

- Note constant range of muons
- Muons have constant kinetic energy!
- Produced in a 2-body decay of the pion





# Consequences

- Cheap, fast-to-start method to study particles
- Quick adoption by university groups
- Detailed cosmic ray studies
- Particles with strangeness (V-shaped tracks)
  - Kaons, hyperons (sigmas, lambdas, xis)
- Gell-Mann et al use the data for their models
- Still used today! E.g., medicine

# References

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